

MAGNETOHYDRODYNAMIC FLOW AND HEAT TRANSFER OF NON-NEWTONIAN POWER-LAW NANOFLUID OVER A ROTATING DISK WITH HALL CURRENT

by

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This work studies the flow and heat transfer of a power-law nanofluid in the presence of an axial uniform magnetic field in the vicinity of a constantly rotating infinite disk. The Hall current effect is taken into consideration. The governing momentum and energy equations are solved numerically by the shooting method. Some of the results obtained for a special case of the problem are compared to the results published in a previous work and are found to be in excellent agreement. The effects of the solid fraction ϕ , the magnetic interaction number M , the Hall current m , and the viscosity index n on the velocity and temperature profiles as well as the local skin friction coefficients and the heat transfer rate are shown graphically.

Key words: *Nanofluid, rotating disk, heat transfer, Hall current*

Introduction.

Rotating disk flow has been at the center of a large number of theoretical and experimental studies in recent years. This is mainly due to its many applications in engineering, such as rotating machinery, lubrication, oceanography, computer storage devices, viscometers, and crystal growth processes. The rotating disk problem was first solved by Von Karmam [1]. He showed that the Navier-Stokes equations for steady flow of a Newtonian incompressible fluid due to a disk rotating far from other solid surface can be reduced to a set of ordinary differential equations. These equations can be solved by using an approximate integral method. Cochran [2] obtained asymptotic solutions for the steady hydrodynamic problem formulated by Von Karman. Benton [3] improved the steady state solutions given by Cochran and extended the problem to transient state.

The rotating disk and stability issues were attacked theoretically, numerically and experimentally by many authors amongst many others, such as Hall [4] and Jarrae and Chauve [5]. The influence of an external uniform magnetic field on the flow due to a rotating disk was studied by Attia and Hassan [6]. Also, the effect of uniform suction or injection through a rotating porous disk on the steady hydrodynamic flow has been investigated by some researchers [7-10]. The effect of thermal radiation on the steady laminar convective hydrodynamic flow of a viscous and electrically conducting fluid due to a rotating disk has considered by Anjali Devi and Uma Devi [11]. The study of hydromagnetic flows with the Hall current has important engineering applications in problems of magnetohydrodynamics generators and Hall accelerations as well as in flight magnetohydrodynamics. Hassan and Attia [12] investigated the steady magnetohydrodynamics boundary layer flow due to an infinite disk rotating with uniform angular velocity in the presence of an axial magnetic field. They neglected induced magnetic field but considered the Hall current and accordingly solved steady state equation numerically using a finite difference approximation. The effect of Hall current on steady laminar convective hydrodynamic flow of an electrically conducting fluid over a porous rotating infinite disk has been examined by Anjali Devi and Uma Devi [13]. Recently, Khidir [14] investigated the effects of viscous dissipation and Ohmic heating on steady MHD convective flow due to a porous rotating disk with variable properties. A number of studies have been reported in the literature focusing on nanofluids because of their industrial and engineering applications, such as electronics, transportation as well as nuclear reactors and biomedicine. Nanofluids can be also used in various biomedical applications

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like cancer therapeutics, nano-drug delivery, nanocryosurgery and cryopreservation. Choi [15] was the first to introduce the term nanofluid to represent a fluid in which nano-scale particles (nano-particles) are suspended in a base fluid with a low thermal conductivity such as water, ethylene glycol and oils [16-17]. The concept of nano-fluids, in recent years, has been proposed as a route for surpassing the performance of heat transfer rate in regular fluids. Experimental studies show that even with a small volumetric fraction of nano-particles (usually less than 5%), the thermal conductivity of the base fluid is enhanced by 10-50% with a remarkable improvement in the convective heat transfer coefficient (see [18-23]). Bachok et al.[24] investigated the steady flow of an incompressible viscous fluid due to a porous rotating disk in a nanofluid via the Keller-box method. They considered two models for the effective thermal conductivity of the nanofluid. They also reported different behaviors for the heat transfer rate at the surface for these methods. Recently, the flow and heat transfer characteristic over a rotating disk immersed in five distinct nanofluids have been investigated by Mustafa [25].

One type of non-Newtonian fluid is the power-law fluid. These fluids are characterized by the property that during its motion the stress is a nonlinear function of the rate of strain. Many of the inelastic non-Newtonian fluids encountered in chemical engineering processes, are known to follow the empirical Ostwald-de Waele model or the so-called "power-law model". In this model the shear stress varies according to a power function of the strain rate. Acrivos et al. [26] investigate the flow of a non-Newtonian fluid (power-law fluid) on a rotating disk. Numerical solution for the flow caused by a disk rotating in liquids with a shear dependent viscosity was first obtained by Mistschka and Ulbricht [27]. Wichterle and Mitschka [28] revised the same study, with a focus on a shear of liquid particles to fit the application of micro-mixing technology. Andersson et al. [29] reconsidered the problem of Mistschka and Ulbricht to test the reliability of their numerical solutions when considering shear-thickening fluids beyond those considered by Mitschka and Ulbricht. The influence of an external magnetic field on the flow due to a rotating disk was studied by several authors [30-32]. Some interesting effects of magnetic field on the flow of a power-law fluid over a rotating disk was examined by Andersson [33]. In all these investigations, the effects of the Hall current were not considered. Recently, the effects of Hall currents on hydromagnetic flow due to a rotating disk have been studied by Attia et al.[34], Hassan et al. [35] and Abdul Maleque et al. [36]. Abo-Eldahab and Salem [37] considered the MHD flow and heat transfer of non-Newtonian power-law fluid with diffusion and chemical reaction on a moving cylinder. Chunying et al. [38] studied the steady flow and heat transfer of a viscous incompressible power-law fluid over a rotating disk.

The aim of the present work is to study the flow and heat transfer due to a rotating disk immersed in a power-law nanofluid in the presence of Hall current using a nanofluid model proposed by Tiwari and Das [39]. The system of nonlinear partial differential equations is transformed into coupled nonlinear ordinary differential equations by means of Von Karman similarity variables and solved numerically by Rung-Kutta method coupled with Shooting technique[40]. The radial, tangential, axial velocities and temperature profiles are sketched for different values of solid volume fraction parameter, power-law index, magnetic field and Hall parameter. The effects of these parameters are discussed.

Formulation of the problem

We consider a steady, laminar, axi-symmetric flow of an incompressible non-Newtonian nanofluid driven solely by an infinite rotating disk. As shown in Fig. 1. Let (r, φ, z) be the set of cylindrical coordinates and let the disk rotate with constant angular velocity Ω and be placed at $z=0$. The surface of the disk is maintained at a uniform temperature T_w , while the temperature of the free stream is T_∞ . An external strong magnetic field is applied in the positive z -direction and has flux density B_0 .

The generalized Ohm's law including Hall current is given in the form [41].

$$\mathbf{J} = \sigma (\mathbf{E} + \mathbf{V} \times \mathbf{B} - \frac{1}{en_e} (\mathbf{J} \times \mathbf{B})) \quad (1)$$

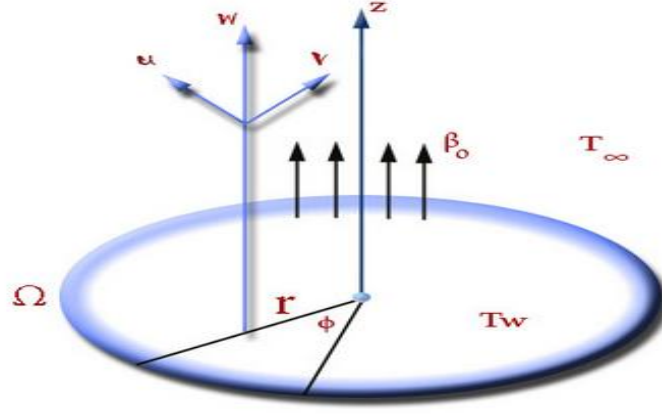


Fig. 1. Physical model and coordinate system

where $\mathbf{V}=u \mathbf{e}_r+v \mathbf{e}_\phi+w \mathbf{e}_z$ is the velocity vector, $\mathbf{B}=(0,0,B_0)$ is the magnetic induction vector, \mathbf{E} is the electric field vector, \mathbf{J} is the current density vector, $\sigma=(e^2 n_e t_e/m_e)$ is the electrical conductivity, t_e is the electron collision time, e is the electron charge, n_e is the electron number density, and m_e is the mass of the electron. In the low-magnetic- Reynolds-number approximation, the induced magnetic field can be ignored. Since no applied or polarization voltage is imposed on the flow field, the electric field vector \mathbf{E} is equal to zero. For large values of the magnetic field strength, the generalized Ohm's law given by Eq.(1) in the absence of an electric field can be solved for \mathbf{J} to yield

$$J_r = \frac{\sigma B_0}{1+m^2}(mv - u), J_\phi = \frac{\sigma B_0}{1+m^2}(v + mu) \quad (2)$$

where $m = w_e t_e$ is the Hall parameter with $w_e (= eB_0/m_e)$ as the electron frequency. It is assumed that the base fluid and the nanoparticles are in thermal equilibrium and no slip occurs between them. The thermophysical properties of the fluid and nanoparticles are given in Table 1. [22].

Table 1. Thermophysical properties of the base fluid and the nanoparticles

Physical properties	Fluid phase(water)	Cu
c_p (J/kgK)	4179	385
ρ (Kg/m ³)	997.1	8933
k (W/mK)	0.613	400
$\alpha \times 10^7$ (m ² /s)	1.47	1163.1

Under these assumptions with the usual boundary layer approximation and using the nanofluid model proposed by Tiwari and Das [39], the governing equations for the conservation of mass, momentum and energy for the problem under consideration (in the presence of Hall current) can be written as

$$\frac{\partial u}{\partial r} + \frac{u}{r} + \frac{\partial w}{\partial z} = 0 \quad (3)$$

$$u \frac{\partial u}{\partial r} - \frac{v^2}{r} + w \frac{\partial u}{\partial z} = -\frac{1}{\rho_{nf}} \frac{\partial p}{\partial r} + \frac{1}{\rho_{nf}} \frac{\partial}{\partial z} \left\{ \mu_{nf} \left[\left(\frac{\partial u}{\partial z} \right)^2 + \left(\frac{\partial v}{\partial z} \right)^2 \right] \frac{n-1}{2} \frac{\partial u}{\partial z} \right\} - \frac{\sigma B_0^2}{\rho_{nf} (1+m^2)} (u - mv) \quad (4)$$

$$u \frac{\partial v}{\partial r} + \frac{uv}{r} + w \frac{\partial v}{\partial z} = \frac{1}{\rho_{nf}} \frac{\partial}{\partial z} \left\{ \mu_{nf} \left[\left(\frac{\partial u}{\partial z} \right)^2 + \left(\frac{\partial v}{\partial z} \right)^2 \right] \frac{n-1}{2} \frac{\partial v}{\partial z} \right\} - \frac{\sigma B_0^2}{\rho_{nf} (1+m^2)} (v + mu) \quad (5)$$

$$u \frac{\partial T}{\partial r} + w \frac{\partial T}{\partial z} = \frac{1}{\rho_{nf}} \frac{\partial}{\partial z} \left\{ k_{nf} \left[\left(\frac{\partial u}{\partial z} \right)^2 + \left(\frac{\partial v}{\partial z} \right)^2 \right]^{\frac{n-1}{2}} \frac{\partial T}{\partial z} \right\} \quad (6)$$

$$-\frac{\partial p}{\partial z} = 0 \quad (7)$$

The boundary conditions are

$$\begin{aligned} u = w = 0, \quad v = \Omega r, \quad T = T_w \quad \text{at} \quad z = 0 \\ u = v = 0, \quad T = T_\infty \quad \text{as} \quad z \rightarrow \infty \end{aligned} \quad (8)$$

Where u , v and w are velocity components in the directions of increasing r , ϕ , z respectively.

$\mu = \mu_{nf} \left[\left(\frac{\partial u}{\partial z} \right)^2 + \left(\frac{\partial v}{\partial z} \right)^2 \right]^{\frac{n-1}{2}}$ is the effective viscosity of nanoliquid, μ_{nf} is the modified consistency

viscosity coefficient, ρ_{nf} is the effective density of nanoliquid and n is the flow behavior index.

The fluid is Newtonian for $n=1$. The fluid are termed pseudoplastic (or shear thinning) for $n<1$ and dilatant (or shear thickening) for $n>1$. Following Zheng et al. [42] we assume the thermal conductivity

is power-law dependent on velocity as $k = k_{nf} \left[\left(\frac{\partial u}{\partial z} \right)^2 + \left(\frac{\partial v}{\partial z} \right)^2 \right]^{\frac{n-1}{2}}$, where k is the effective thermal

conductivity of nanoliquid and k_{nf} is the modified thermal conductivity of the nanoliquid. The properties of nanofluids are defined as follows [17]:

$$\begin{aligned} \rho_{nf} &= (1-\phi)\rho_f + \phi\rho_s, \quad \mu_{nf} = \frac{\mu_f}{(1-\phi)^{2.5}}, \quad \alpha_{nf} = \frac{k_{nf}}{(\rho c_p)_{nf}} \\ (\rho c_p)_{nf} &= (1-\phi)(\rho c_p)_f + \phi(\rho c_p)_s, \quad \frac{k_{nf}}{k_f} = \frac{(k_s + 2k_f) - 2\phi(k_f - k_s)}{(k_s + 2k_f) + 2\phi(k_f - k_s)} \end{aligned} \quad (9)$$

where ϕ is the solid volume fraction of nanoliquid, μ_f is the viscosity of the basic fluid, ρ_s is the density of the solid, $(\rho c_p)_f$ heat capacity of the base fluid, $(\rho c_p)_s$ heat capacity of solid, k_f thermal conductivity of base fluid and k_s thermal conductivity of solid. By introducing the Von Karaman transformation [1], one finds that

$$\begin{aligned} u &= r \Omega F(\eta), \quad v = r \Omega G(\eta) \\ w &= \left(\frac{\Omega^{1-2n}}{k/\rho} \right)^{-1/(n+1)} r^{(n-1)/(n+1)} H(\eta), \quad \theta(\eta) = \frac{T - T_\infty}{T_w - T_\infty} \end{aligned} \quad (10)$$

where

$$\eta = z \left(\frac{\Omega^{2-n}}{k/\rho} \right)^{1/(n+1)} r^{(1-n)/(1+n)} \quad (11)$$

is the generalized dimensionless similarity variable originally proposed by Mitschka [43] and F , G and H are non-dimensional function of η .

In the present study the magnetic interaction number M was investigated to be $M = \frac{\sigma B_o^2}{\rho \Omega}$

which represents the ratio between the magnetic force to the fluid inertia force. With these definitions, Eqs (3)-(6), Eq. (8) take the form

$$H' + 2F + F' \frac{1-n}{1+n} \eta = 0 \quad (12)$$

$$\frac{1}{(1-\phi)^{2.5}(1-\phi+\phi\frac{\rho_s}{\rho_f})} \frac{d}{d\eta} [(F'^2 + G'^2)^{\frac{n-1}{2}} F\eta] - (H + \frac{1-n}{1+n} F\eta) F' - F^2 + G^2 - \frac{M}{1+m^2} (F - mG) = 0 \quad (13)$$

$$\frac{1}{(1-\phi)^{2.5}(1-\phi+\phi\frac{\rho_s}{\rho_f})} \frac{d}{d\eta} [(F'^2 + G'^2)^{\frac{n-1}{2}} G\eta] - (H + \frac{1-n}{1+n} F\eta) G' - 2FG - \frac{M}{1+m^2} (mF + G) = 0 \quad (14)$$

$$\frac{1}{Pr} \frac{k_{nf}/k_f}{1-\phi+\phi\frac{(\rho c_p)_s}{(\rho c_p)_f}} \frac{d}{d\eta} [(F'^2 + G'^2)^{\frac{n-1}{2}} \theta\eta] - (H + \frac{1-n}{1+n} F\eta) \theta' = 0 \quad (15)$$

Where the prime denote the differentiation with respect to similarity variable η

Boundary conditions are

$$F=0, G=1, H=0 \quad \theta = 1 \quad \text{at} \quad \eta = 0$$

$$F=0, G=0, \theta = 0 \quad \text{as} \quad \eta \rightarrow \infty \quad (16)$$

Results and discussion

The system of transformed equations (12)-(15) with the corresponding boundary conditions (16) have been solved numerically by means of the fourth-order Runge-Kutta method with systematic estimates of $F'(0)$, $G'(0)$ and $\theta'(0)$ by shooting technique. In the present calculations, step size of $\Delta\eta = 0.001$ and $\eta_{\max} = 15$ were found to be satisfactory in obtaining sufficient accuracy within a tolerance less than 10^{-8} in nearly all cases. Comparison is made with the available published data in Table 2 for $F'(0)$, $G'(0)$ and $-H(\infty)$ which show a good agreement with the present results.

Table 2. Comparison of some of the present data with results of Anderson [33] for the particular case of a Newtonian fluid (n=1) and for m=0

M	F'(0)		-G'(0)		-H(∞)	
	Anderson [33]	present	Anderson [33]	present	Anderson [33]	Present
0	0.5101	0.51023254	0.6160	0.61592206	0.8827	0.88447335
0.5	0.3851	0.38513252	0.8487	0.84872382	0.4589	0.45888794
1	0.3093	0.30925791	1.0691	1.06905343	0.2533	0.25331375
2	0.2306	0.23055903	1.4421	1.44209404	0.1086	0.10858343
4	0.1657	0.16570305	2.0103	2.01026675	0.0408	0.04077522

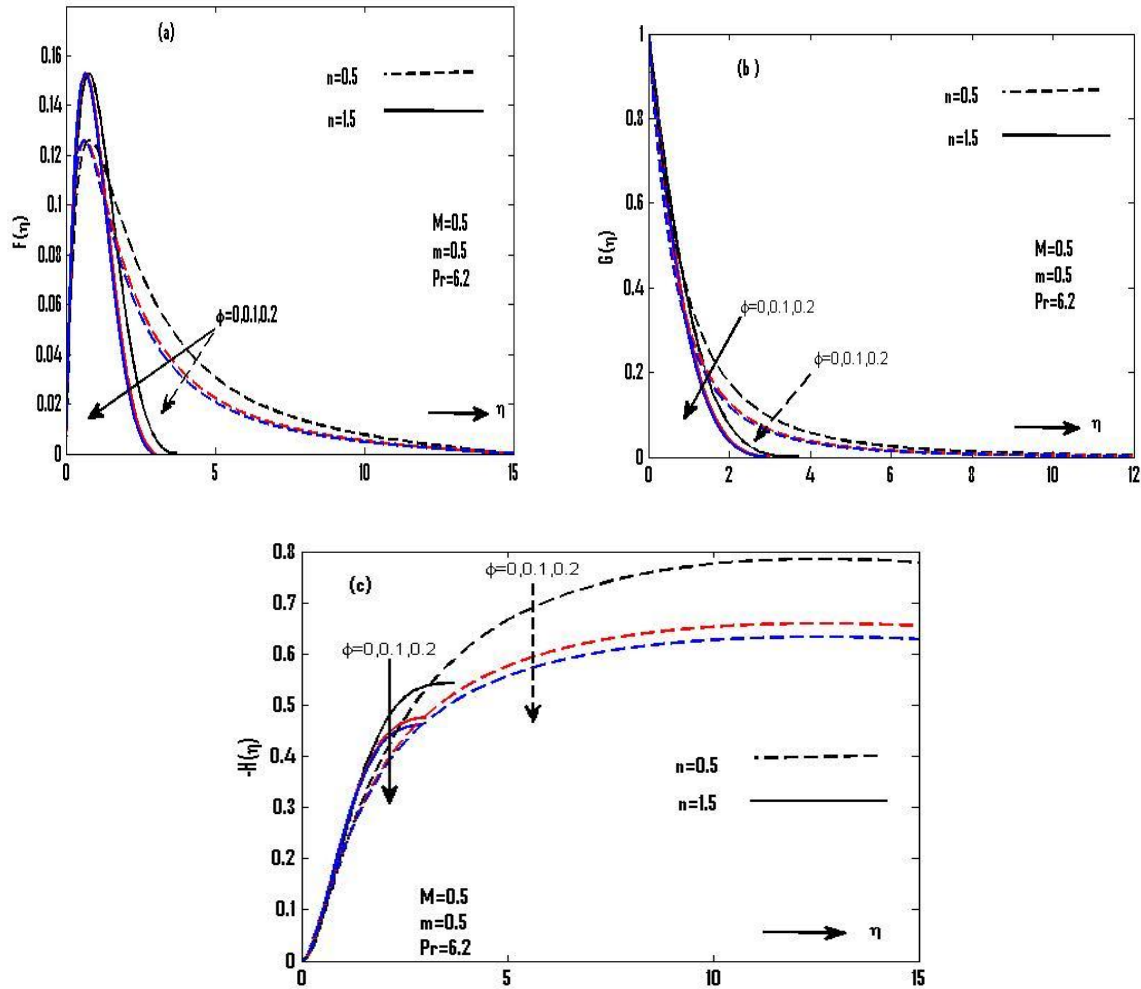


Fig. 2. Effect of solid volume fraction on (a) radial velocity profiles , (b) tangential velocity, Profiles (c) axial velocity profiles.

The effects of solid volume fraction ϕ on the velocity profiles in the radial, tangential and axial directions, F , G and $-H$ for a shear-thinning fluid ($n < 1$) and a shear-thickening fluid ($n > 1$) are shown in Fig. 2(a)-(c). Due to the existence of the centrifugal force, the radial velocity attains a maximum value close to the surface of the disk for all values of ϕ . With increasing the values of solid volume fraction parameter, radial velocity increases near the surface of the disk, but towards the end it decreases. Also, the rate of its convergence to its limiting value (i.e. zero) is faster for larger value of ϕ and n . In Figs. 2(a) and (b), we observe that, far from the surface of the disk, the radial and tangential velocity profiles decreases as n increases. The reason for such behavior is that increasing the power-law index n tends to increase the viscous forces and decelerate the flow, thus decreasing the velocity profiles. Moreover, for $n=1.5$, the axial flow (Fig. 2(c)) fails to approach an asymptotic limit for large η in a regular manner as two other velocity components, F and G , tend to zero at the edge of the boundary layer. This observation can easily be ascribed to an anticipated deterioration of the numerical accuracy caused by the thickening of the boundary layer.

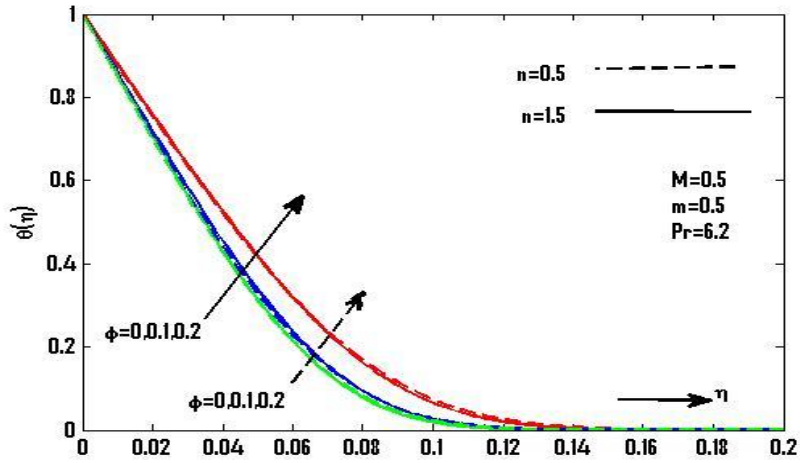


Fig. 3. Effect of solid volume fraction on the temperature profiles.

Fig. 3 shows the influence of solid volume fraction ϕ on the temperature distributions for two different values of the power-law index n . The dimensionless temperature profiles $\theta(\eta)$ decrease monotonically from T_m (i.e. $\theta = 1$) at the surface of the disk to T_∞ (i.e. $\theta = 0$) at infinity. With increase in the solid volume fraction ϕ , the temperature increases across the boundary layer for both values of the power-law index $n < 1$ and $n > 1$. Also, the temperature distribution across the boundary layer slightly increases with increase in values of n due to the little enhancement of the viscous forces, but the effect is opposite far away from the surface of the disk.

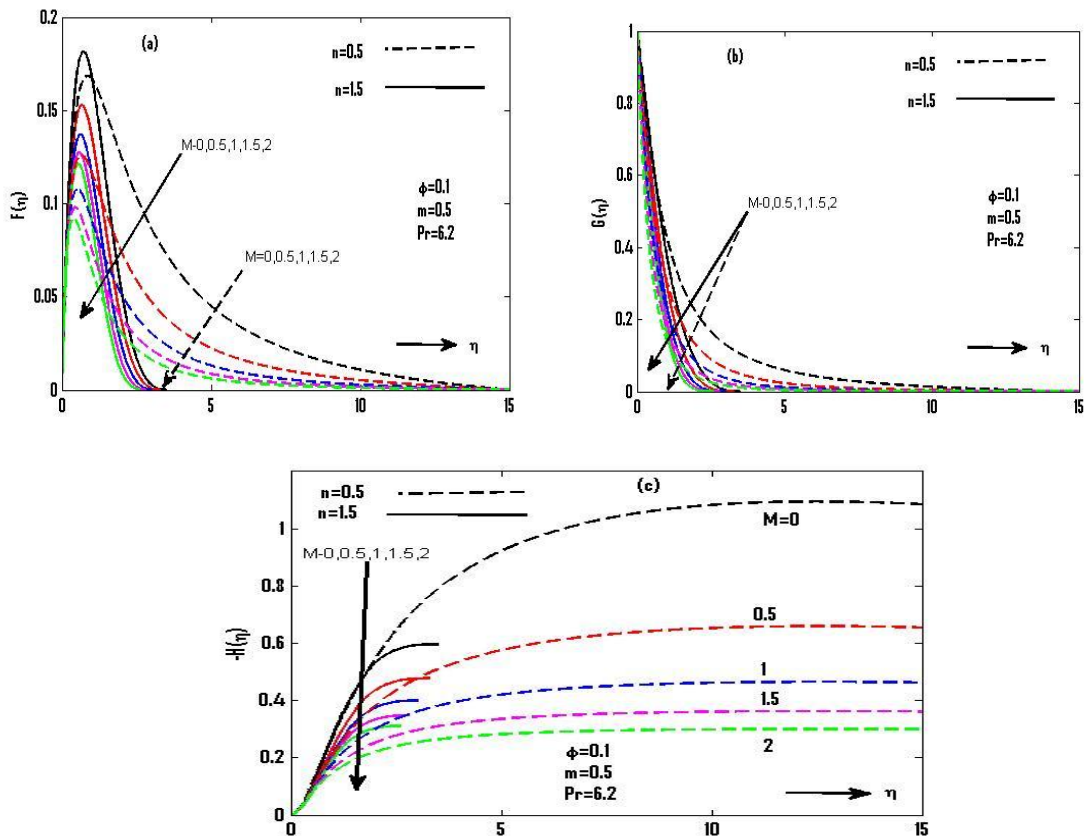


Fig. 4. Effect of magnetic field on (a) radial velocity profiles, (b) tangential velocity, Profiles (c) axial velocity profiles.

Figs 4(a)-(c) depicts the variation of velocity profiles in the radial, tangential, and axial distributions, F , G and H for different values of n and M . Application of a magnetic field produces a resistive force called the Lorentz force. This force has a tendency to slow the flow around the disk. This is depicted by the decreases in the radial, tangential, and axial velocity profiles as M increases, as shown in Figs. 4(a)-(c). The magnetic field effect is also observed to be more pronounced at lower values of M and n because a nanofluid with a thinning boundary layer is more susceptible to magnetic force effects. Also, for a large value of n ($n=1.5$), the axial flow (Fig. 4(c)) fails to approach an asymptotic limit for large η in a regular manner as two other velocity components, F and G , tends to zero at the edge of the boundary layer.

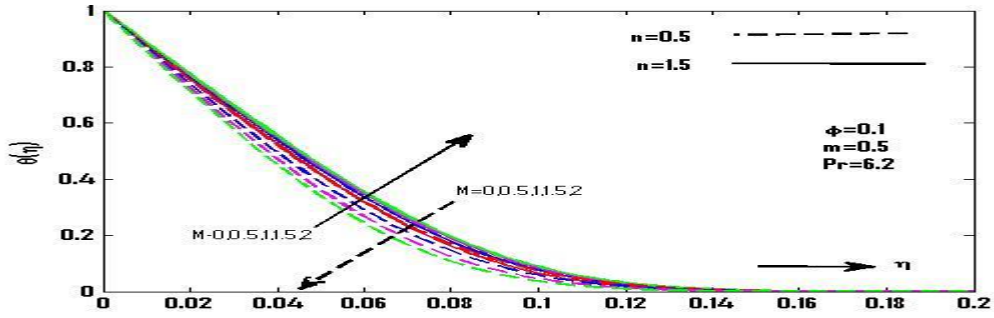


Fig. 5. Effect of magnetic field on the temperature profiles.

Fig. 5. reveals the influence of magnetic field M on the temperature profile for pseudo-plastic and dilatants nanofluids. The temperature profile may be enhanced or reduced by increasing M , depending on the fluid index n . In this figure it is observed that the temperature decreases with increasing M for pseudo-plastic nanofluid ($n < 1$), but increases with M for dilatant nanofluid ($n > 1$). Also, the influence of the magnetic field M is more obvious for a pseudo-plastic nanofluid than for a dilatants nanofluid.

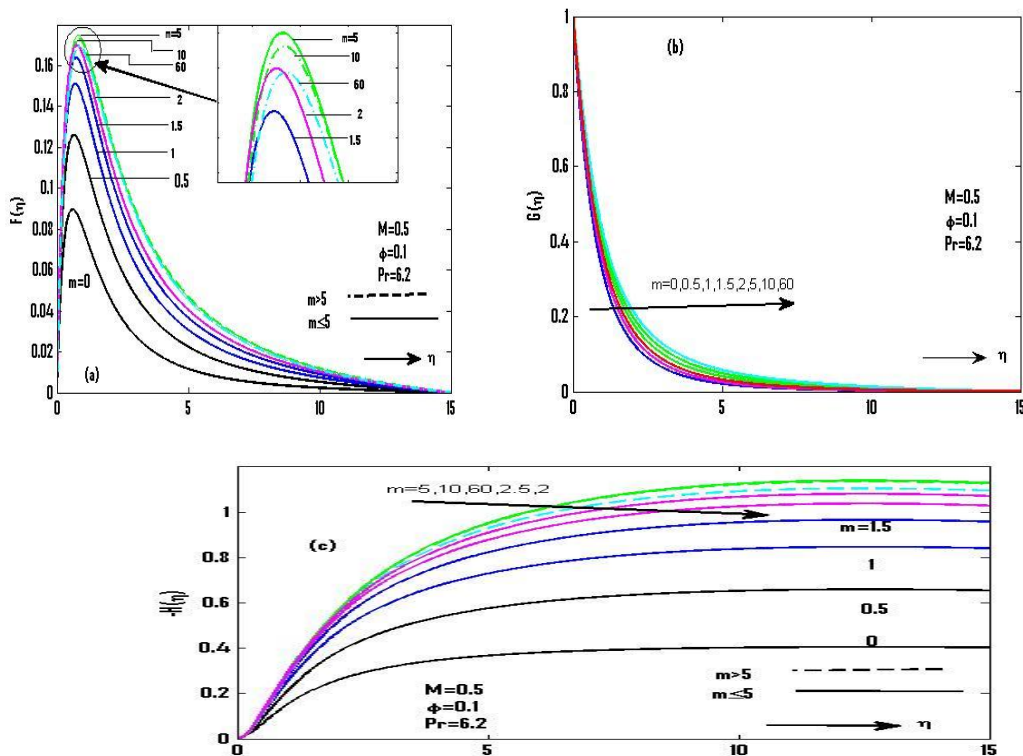


Fig. 6. Effect of Hall parameter on (a) radial velocity profiles , (b) tangential velocity, Profiles (c) axial velocity profiles

Figs. 6(a)-(c) and 6 show the radial, tangential, axial velocities and temperature at different values of the Hall current m for pseudo-plastic nanofluid. The radial and axial velocity profiles increase due to the increase in the magnitude of m in the interval $0 \leq m \leq 5$, Figs. 6(a) and (c). However, beyond $m=5$, the radial and axial velocity profiles decrease as the Hall parameter m increases. Furthermore, this decrease in F and $-H$ is almost negligible for large m ($m > 60$). The phenomenon for small and large values of the Hall parameter m has been explained by Hassan and Attia [34]. The tangential velocity profiles G , as shown in Fig. 6(b), are found to increase for all values of Hall parameter m . Fig. 7 shows that the temperature profile decreases for $m \leq 1.5$ and increase for $m > 1.5$. This increase in θ is almost negligible for large values of m .

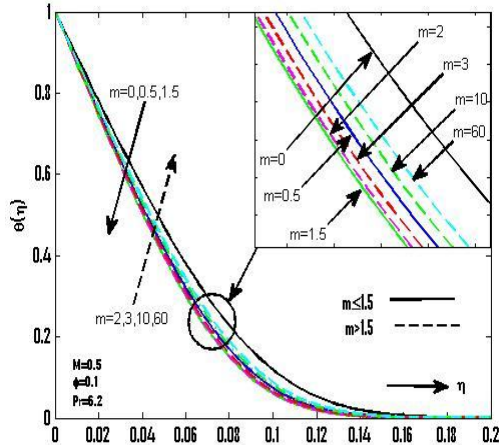


Fig. 7. Effect of Hall parameter on the temperature profiles.

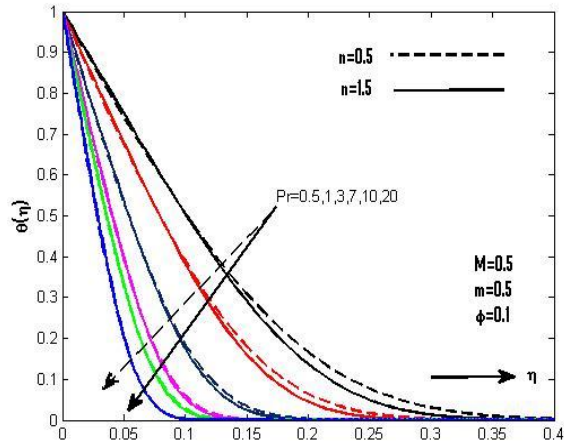


Fig. 8. Effect of Prandtl number on the temperature profiles

Fig. 8 shows the usual effect of Prandtl number Pr on the temperature profiles for pseudo-plastic and dilatants nanofluids. It can be seen that the Prandtl number decreases the temperature distribution throughout the boundary layer. This is due to the fact that there would be a decrease of thermal boundary layer thickness with the increase of Prandtl number. In addition, the temperature profile is less affected by the fluid index n for $Pr=20$. This behavior implies that the fluid of smaller Prandtl number is more responsive to fluid index n than the fluid having a larger Prandtl number.

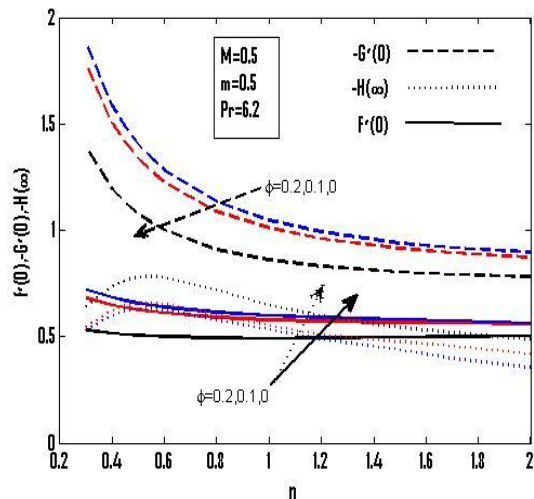


Fig. 9. Effect of solid volume fraction on $F'(0)$, $-G'(0)$, $-H'(\infty)$

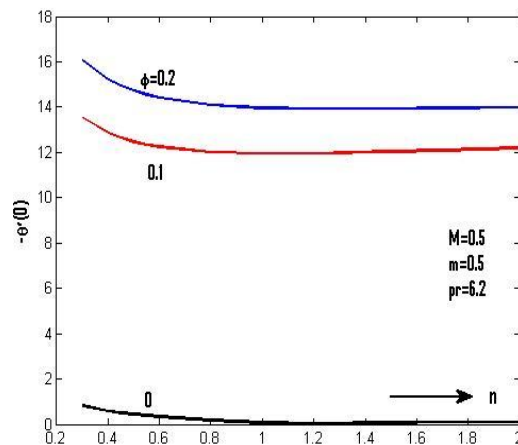


Fig. 10. Effect of solid volume fraction on the rate of heat transfer $-\theta'(0)$

Finally, the values of the radial and tangential skin frictions $F'(0)$ and $G'(0)$ at the disk surface, as well as the axial inflow $-H(\infty)$ and heat flux on the surface $-\theta'(0)$ are plotted against the power-law index n in Figs. 9 and 10. We observe that an increase in the value of the solid volume

fraction parameter ϕ may lead to a decrease in the radial skin friction $F'(0)$ and the axial inflow $-H(\infty)$, yet it increases the tangential skin friction coefficient $G'(0)$ and rate of heat transfer $-\theta'(0)$. It is also to note that the radial skin Friction decreases with n for various values of the solid volume fraction ϕ ; However the trend is reversed for large ϕ , especially when the fluid is dilatant nanofluid ($n > 1$).

Conclusions

The present work deals with the magnetohydrodynamic flow and heat transfer of non-Newtonian power-law nanofluid due to a rotating disk. The governing fundamental equations are approximated by a system of nonlinear ordinary differential equations by using a similarity transformation, and are solved numerically by using the Rung-Kutta and shooting methods. The numerical results obtained agree very well with the previously published data in some particular cases of the present study. The effects of the solid volume fraction ϕ and the magnetic interaction number M on nanofluid velocity and temperature are discussed for shear-thinning and shear-thickening non-Newtonian nanofluids. The highlights of this study are:

- The solid volume fraction decreases the magnitude of the axial and tangential nanofluid velocity components and increases the magnitude of the radial nanofluid velocity component and temperature profiles throughout the boundary layer.
- The obtained numerical results show that the Hall parameter m has an interesting effect on the radial and axial nanofluid velocity profiles. For large values of $m (> 5)$, the resistive effects of the magnetic field are diminished, hence, the radial and axial nanofluid velocity profiles decrease with the increase of m while the tangential nanofluid velocity profile increases for all values of m .
- The temperature profile decreases with increasing m in the range $0 \leq m \leq 1.5$, when for $m > 1.5$, the temperature profile increases with increasing m .
- The magnetic field has increasing effects on the temperature profiles for dilatant non-Newtonian nanofluid, whereas the opposite is observed for pseudo-plastic non-Newtonian nanofluid.

Nomenclature

B	external uniform magnetic field ($\text{kg/s}^2\text{A}$)	η	a scale boundary-layer coordinate
B_o	constant magnetic flux density ($\text{kg/s}^2\text{A}$)	θ	self-similar temperature
c_p	specific heat at constant pressure (J/kg k)	μ	dynamic viscosity (Ns/m^2)
E	electric field vector	ν	kinematic viscosity (m^2/s)
e	charge of electron	ρ	density (kg/m^3)
f	self-similar velocity	σ	electrical conductivity (s/m)
g	acceleration due to gravity (m/s^2)	ϕ	nanoparticle volume fraction
J	current density	<i>Subscripts</i>	
k	thermal conductivity	f	fluid phase
m	Hall parameter	s	solid phase
n_e	number density of electrons	w	condition of the wall
T	fluid temperature (k)	∞	ambient condition
t_e	electron collision time		
u	radial velocity		
v	tangential velocity		
w	axial velocity		
(r, ϕ, z)	cylindrical polar coordinates		

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